

Development & Extratropical Applications of the CYGNSS Surface Heat Flux Product 回社注回 Juan A. Crespo*, Shakeel Asharaf, Catherine M. Naud, Rosa Luna-Niño, James Booth, & Derek J. Posselt *UCLA Joint Institute for Regional Earth System Science and Engineering | Jet Propulsion Laboratory, California Institute of Technology *Contact: Juan.A.Crespo@jpl.nasa.gov

Introduction

- Surface latent (LHF) and sensible (SHF) heat fluxes aid in the transport of heat and energy between the atmosphere and ocean
 - Driven by winds and air-sea temperature/humidity differences
- LHF and SHF increase baroclinicity and instability within the boundary layer, influencing climate/weather systems like:
 - Tropical & Extratropical Cyclones (TCs/ETCs)
 - **Atmospheric Rivers (ARs)**
 - Tropical Convection (e.g. MJO)



- CYGNSS's dense coverage over tropical/subtropical oceans has been used to estimate surface heat fluxes at higher frequencies
 - A combination of other datasets, like reanalysis for temperature & humidity, are used to estimate LHF & SHF

Scan to access

CDR V1.2

from the

PO.DAAC



CYGNSS Flux Algorithm

- Input CYGNSS L2 Wind Speed observations (FDS & YSLF) and ERA5 Reanalysis for temperature and humidity
 - ERA5 is co-located to CYGNSS specular points with a tri-linear interpolation in time and space
 - Previous used MERRA-2 for thermodynamic variables
- Coupled Ocean-Atmosphere Response Experiment (COARE), Version 3.5 is used to estimate LHF and SHF
 - Parameterizes surface heat flux drag coefficients (C_D) as function of gustiness, surface roughness, & atmospheric stability
 - Verified for wind speeds up to 25 m/s (QC flags for wind over 25 m/s)
 - Assume equivalent neutral winds from CYGNSS in upcoming version (SDR V3.2) & make necessary corrections to COARE algorithm
- Output as a Level 2 product with the same number of specular points as their respective Level 2 wind speed product FDS and YSLF ocean surface heat flux estimates

LHF

36.82

18.04

32.09

0.79

54,989

RMSD

σ

SHF

9.07

5.52

7.19

0.58

54,989

Current CYGNSS Flux Products and Validation



Fig. 1: Mean CYGNSS LHF observations (CDR V1.2) of Dec., Jan., Feb. (DJF) 2018-2023. Units: W/m²

CYGNSS L2 SDR V3.1 Winds & ERA5 Reanalysis

CYGNSS L2 CDR V1.2 Winds & ER5 Reanalysis

CYGNSS L2 SDR V3.2 Winds & ERA5 Reanalysis

Level 2 SDR V3.2 to be released Summer 2024

Level 2 Science Data Record (SDR) V2.0

Level 2 Climate Data Record (CDR) V1.2

- Comparisons of CYGNSS LHF and SHF estimates to tropical buoys show how well CYGNSS fluxes compare to in-situ observations (Fig. 2)
- CYGNSS fluxes perform well at lower flux values
 - Greater scatter at higher flux values due to wind speed difference (LHF) and Ts-Ta differences (SHF)
- Improvement in surface heat fluxes as the CYGNSS wind speeds improve
 - Improvement in fluxes with corrections to the COARE algorithm to factor in CYGNSS's equivalent neutral wind speed assumptions



Fig. 2: Comparisons of the upcoming CYGNSS SDR V3.2 flux product to tropical buoy data. Table to the left shows statistical values from the graph above.







CYGNSS Extratropical Cyclone Observations

CYGNSS observations show higher LHF values as the ETC develops in WPAC (Fig. 3)



Fig. 3: Left: CYGNSS LHF (W/m²) at 2019-12-27 18z & 12-28 15z; Middle: IMERG Precip rate (mm/hr) at 2019-12-28 & 12-29 at 18z; Right: HYSPLIT forward trajectories from areas of high LHF values observed by CYGNSS.

- HYSPLIT preliminary trajectories show parcels from high flux areas & Omega following the ETC, vertical movement as well
- Parcel locations correlate with strong precipitation ~24 hours later
- Stronger surface heat fluxes when ETC matures
- HYSPLIT trajectories show a similar pattern but do not line up with the precipitation observed
- Precipitation rates 24 hours after LHF maxima weaker than before
- If fluxes play a role, fluxes earlier or before development likely play a larger role than fluxes when the ETC has matured
- Composite analysis of strong and weak fluxes in the area of the Warm Sector (WS) 24 hours prior to arrival (Fig. 4) suggests that fluxes before ETC arrival play a significant role in ETC development.



Top Height) between ETCs with weak and strong LHF values in the Warm Sector 24-hours prior to ETC arrival.



CYGNSS Atmospheric River Observations

- CYGNSS observations show high LHF values associated with developing and matured ARs in the Northern Pacific Ocean
 - Strong LHF values present near Hawai'i associated with a Kona Low (Fig. 5) before developing into the 2019 Valentine's Day AR
 - HYSPLIT trajectories from areas of high LHF show parcels following the path of the AR, lining up with landfall and early precipitation development
 - GPM PR observations show significant mesoscale structure within the AR as it forms, lining up with some of the parcel trajectories
- AR Family in January 2023 impacted the US West Coast, with most of the ARs having strong LHF values (Fig. 6) Trajectories from areas of high fluxes show energy being transported upstream to the other ARs, possibly strengthening them Some parcels (just a degree apart) follow the path of the AR, suggesting that the location and timing of the fluxes are highly important



Fig. 5: CYGNSS LHF (W/m²) at 2019-02-10 & 02-11 at 03z (left) as the Valentine's Day Atmospheric River begins to form. GPM Precipitation Radar observations (right) of the AR on 2019-02-11 at 00z, with HYSPLIT forward trajectories (top) linking them to observed LHF.

Fig. 6: LHF observations (left) from an AR Family in January 2023, with HYPLIT forward trajectories from areas of high LHF.

Conclusions

- CYGNSS Ocean Surface Heat Fluxes have become a valuable tool for observing and understanding air-sea processes over the tropical and subtropical oceans.
- CYGNSS provides valuable flux observations as ETCs and ARs develop in the lower midlatitudes
 - Fluxes earlier in their lifecycles likely play a role in cloud and precipitation development, but when matured, the fluxes (though stronger) do not.
 - Fluxes associated with ARs and ETCs may be transported upstream, strengthening other ETCs/ARs in a family. But more tests are needed.



Future Developments

Future CYGNSS flux products will include Local Solar Time (LST), which will allow us to factor in diurnal cycles in ETCs and ARs on longer time scales (Fig. 7)

Further trajectory analysis is needed through HYSPLIT or LAGRANTO in order to better analyze the link between air-sea interactions and mesoscale cloud/precipitation formation in ETCs and ARs.

Location and timing are important